

# *Atomic and Molecular Physics for Astronomical Observations*

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# *Layout of Talk*

- Overview of SciDAC (Scientific Discovery through Advanced Computing) project to provide atomic data for use by fusion plasma modelers
  - Atomic data provision and formats
  - Electron-impact excitation
  - Electron-impact ionization
  - Dielectronic recombination
  - Other atomic physics calculations (spin-offs!)
- Recent/current atomic scattering calculations relevant to astrophysics
- Some planned projects of relevance
- Discussion of the current atomic data needs of astrophysics

# *Terascale computational atomic physics for controlled fusion energy*



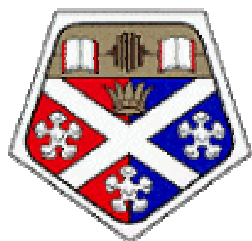
OAK RIDGE NATIONAL LABORATORY



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# *Atomic Data Provision*



- *Controlled Fusion Atomic Data Center at Oak Ridge National Laboratory, TN*
- *<http://www-cfadc.phy.ornl.gov>*



- *Atomic **D**ata and **A**nalysis Structure, UK consortium based at Strathclyde University, Glasgow, UK, and JET, Oxford, UK*
- *<http://adas.phys.strath.ac.uk/adas/docs/manual>*
- *H. P. Summers, ADAS Users Manual (2<sup>nd</sup> Edition) (2003)*

# *Scientific Utilization of Advanced Computing Platforms*

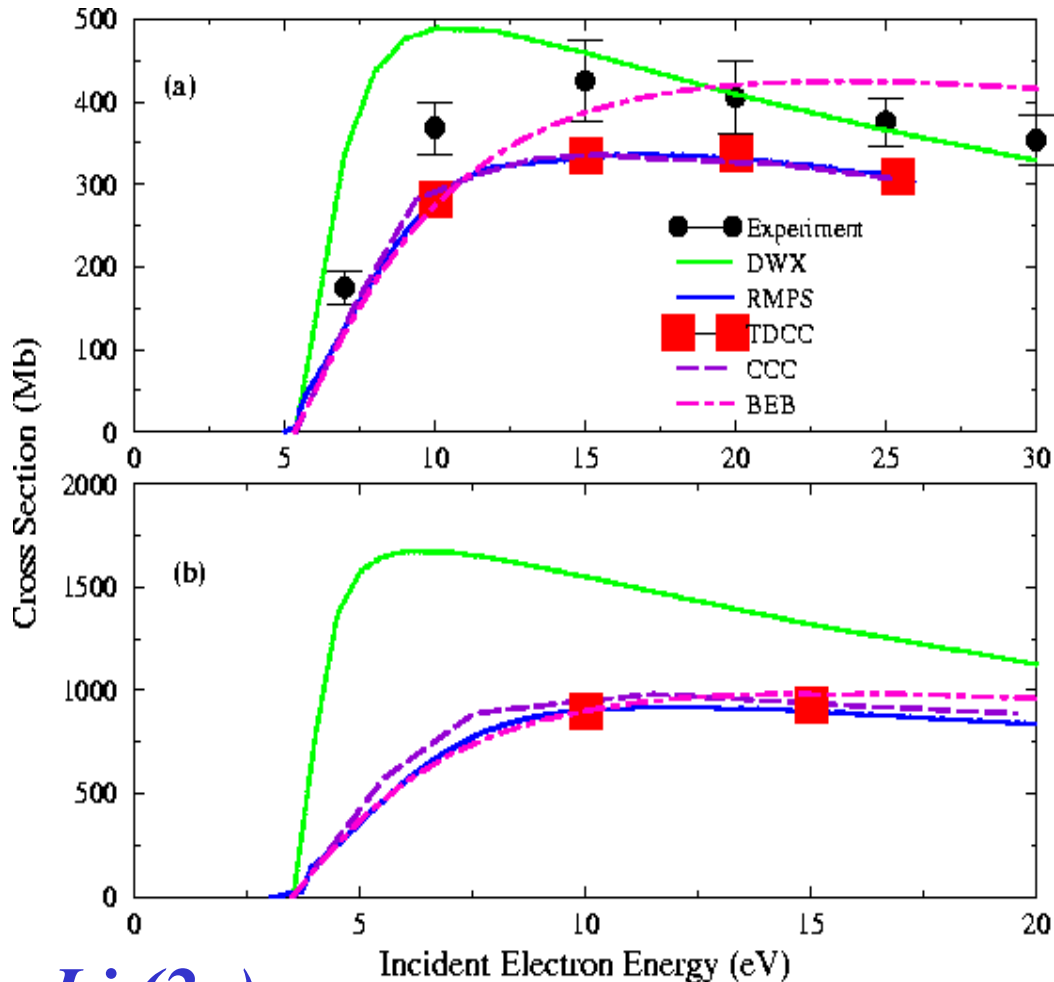
- **Address fundamental problems in atomic and molecular physics**
  - *One-center three-body break-up problem*
  - *Two-center two-body scattering problems*
  - *Collisions in intense external fields*
- **Respond to current needs in controlled fusion energy**
  - *Lithium database for DIII-D plasma transport codes*
  - *Beryllium database for wall component studies at UCSD (PISCES)*
  - *Heavy species database for wall erosion studies at JET*
- **Apply computational methods to other scientific areas**
  - *Simulations of cold Rydberg plasmas*
  - *Time-dependent studies of Bose-Einstein condensates*
  - *Iron spectral analysis of active galactic nuclei*

# *Electron-Impact Ionization Data for modeling fusion plasmas*

- Program initiated several years ago to completely revise electron-impact excitation and ionization data for lithium
- Excitation and ionization of Li, Li<sup>+</sup>, Li<sup>2+</sup> examined
- Significant differences found between our TDCC calculations and experiment for neutral Li
- Three non-perturbative techniques however gave excellent agreement for Li – suggest experimental measurements need revision
- Data placed in format suitable for use by astrophysical (and fusion) plasma modelers
- Will be used to support plasma edge diagnostics made using a Li beam at the DIIID tokamak in San Diego (GA)

# *Electron-impact ionization of lithium*

## *Li (2s)*



## *Li (2p)*

**Demonstrates importance of treating three-body Coulomb problem using non-perturbative methods – large differences with perturbative calculations of 50 – 100 % in ionization cross section**

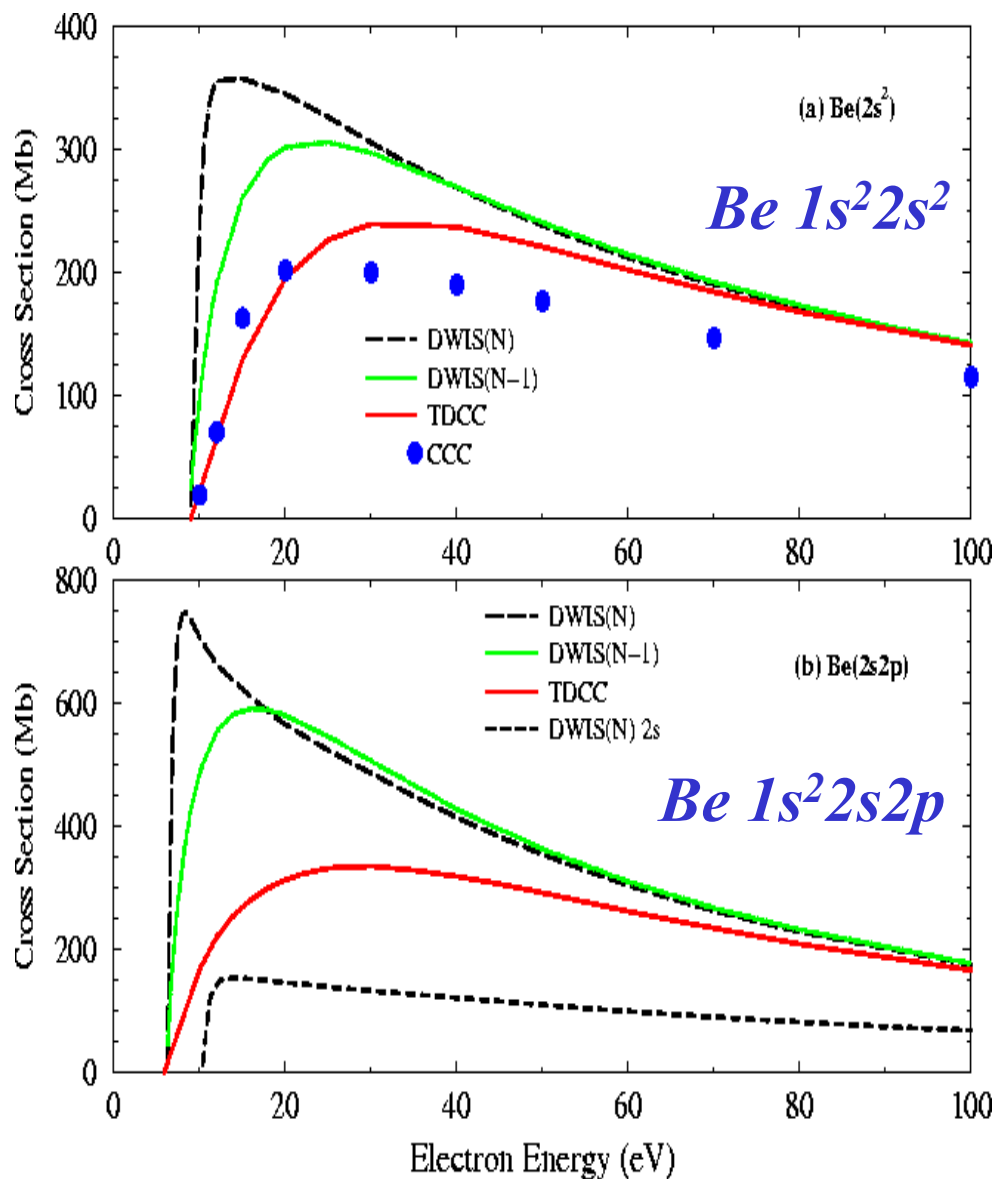
***J. Colgan, M. S. Pindzola, D. M. Mitnik, D. C. Griffin, and I. Bray, Phys. Rev. Letts. 87, 213201 (2001)***

# *Electron-Impact Ionization Data for modeling plasmas*

- Program currently underway to revise electron-impact excitation and ionization data for beryllium
- Recently calculated electron-impact ionization of Be, Be<sup>+</sup>, Be<sup>2+</sup>, and Be<sup>3+</sup>. Excitation calculations well underway
- Few experimental measurements due to toxicity of beryllium
- Significant differences found between TDCC calculations and perturbative distorted-wave calculations for neutral Be
- Differences between non-perturbative and perturbative decrease as ion stage increases
- Data placed in format suitable for use by astrophysical (and fusion) plasma modelers – specifically for the PISCES plasma fusion device at UCSD which will use Be as a potential plasma-facing wall component

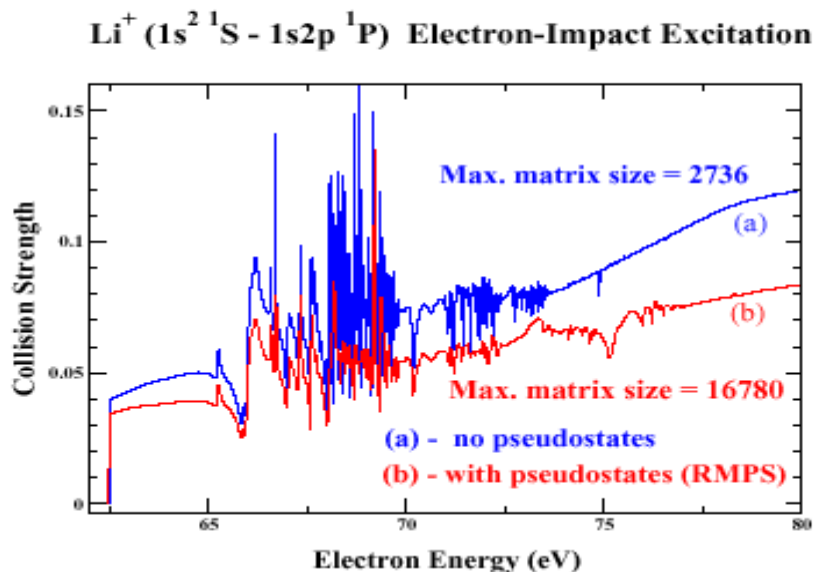
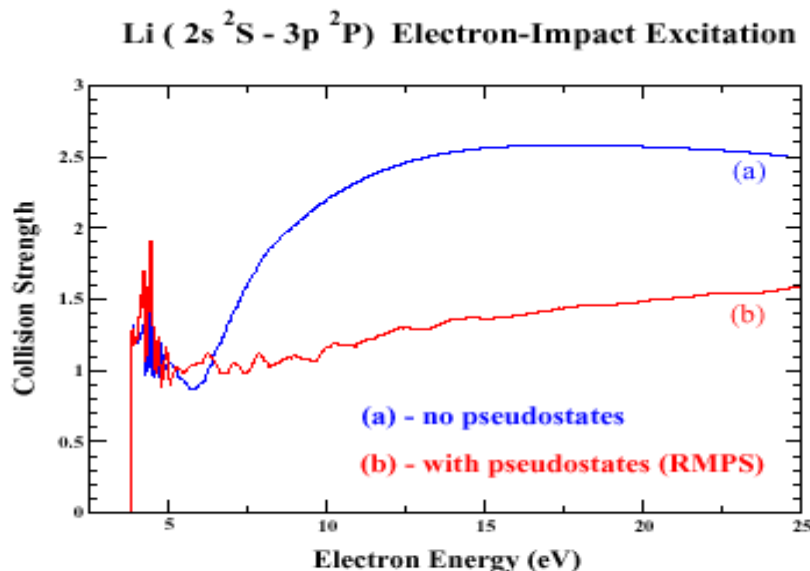


# Electron-impact ionization of Be



- **Red line: TDCC**
- **Green line: DWIS(N)**
- **Black line: DWIS(N-1)**
- **Blue circles: CCC**
- Large differences near ionization threshold and cross section peak between the perturbative and nonperturbative calculations
- No experimental measurements with which to compare for either configuration
- *J. Colgan, S. D. Loch, M. S. Pindzola, C. P. Ballance, and D. C. Griffin, Phys. Rev. A, submitted (2003).*

# *Electron-impact excitation of Li and Li<sup>+</sup>*



- R-matrix with Pseudo States (RMPS) method has been used for many electron-impact excitation cross section calculations:

- In these calculations use of pseudo states reflected the importance of accurately representing the coupling to the continuum

- **Li**: D. C. Griffin et al, Phys. Rev. A **64**, 032718 (2001).

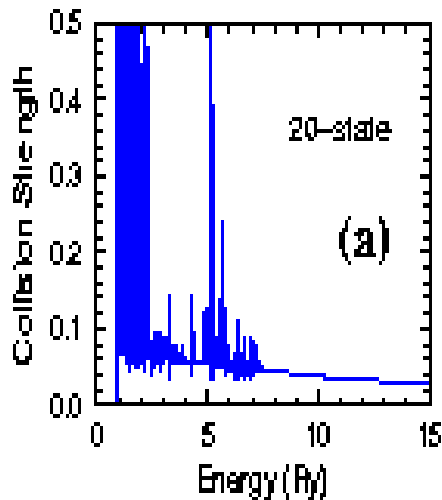
- **Li<sup>+</sup>**: C. P. Ballance et al, J. Phys. B **36**, 235 (2003).

- **C<sup>2+</sup>**: D. M. Mitnik et al, J. Phys. B **36**, 717 (2003).

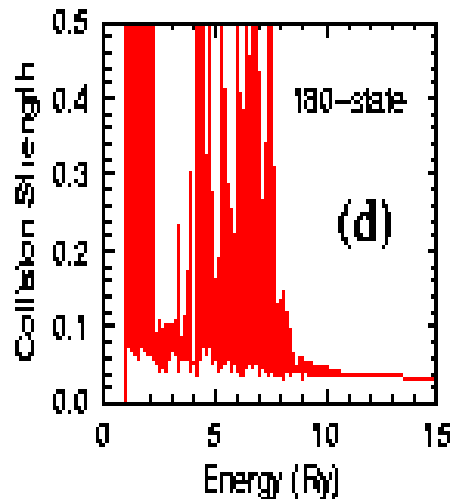
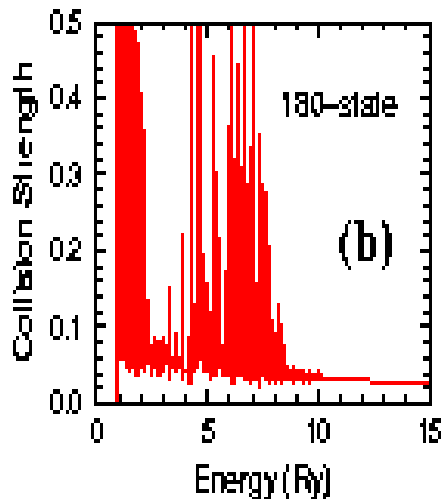
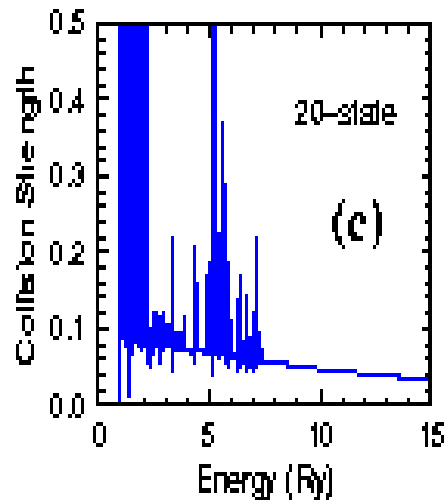
- **C<sup>3+</sup>**: D. C. Griffin et al, J. Phys. B **33**, 1013 (2000).

# *Electron-impact excitation of heavier ions*

$\text{Ne}^3+ (2s^2 2p^2 P_{1,2} - 2s 2p^2 {}^3P_{3,2})$



$\text{Ne}^3+ (2s^2 2p^2 P_{3,2} - 2s 2p^2 {}^3P_{3,2})$



• Large R-matrix calculations (with no pseudo states) contain many levels to represent the low-lying bound states accurately

•  $\text{Ne}^+$ : Griffin et al, J. Phys. B **34**, 4401 (2001).

•  $\text{Ne}^{4+}$ : Griffin et al, J. Phys. B **33**, 4389 (2000).

•  $\text{Ne}^{5+}$ : Mitnik et al, J. Phys. B **34**, 4455 (2001).

•  $\text{Fe}^{20+}$ : Badnell and Griffin, J. Phys. B **34**, 681 (2001).

•  $\text{Fe}^{21+}$ : Badnell et al, J. Phys. B **34**, 5071 (2001).

•  $\text{Fe}^{23+}$ : Whiteford et al, J. Phys. B **35**, 3729 (2002).

•  $\text{Fe}^{24+}$ : Whiteford et al, J. Phys. B **34**, 3179 (2002).

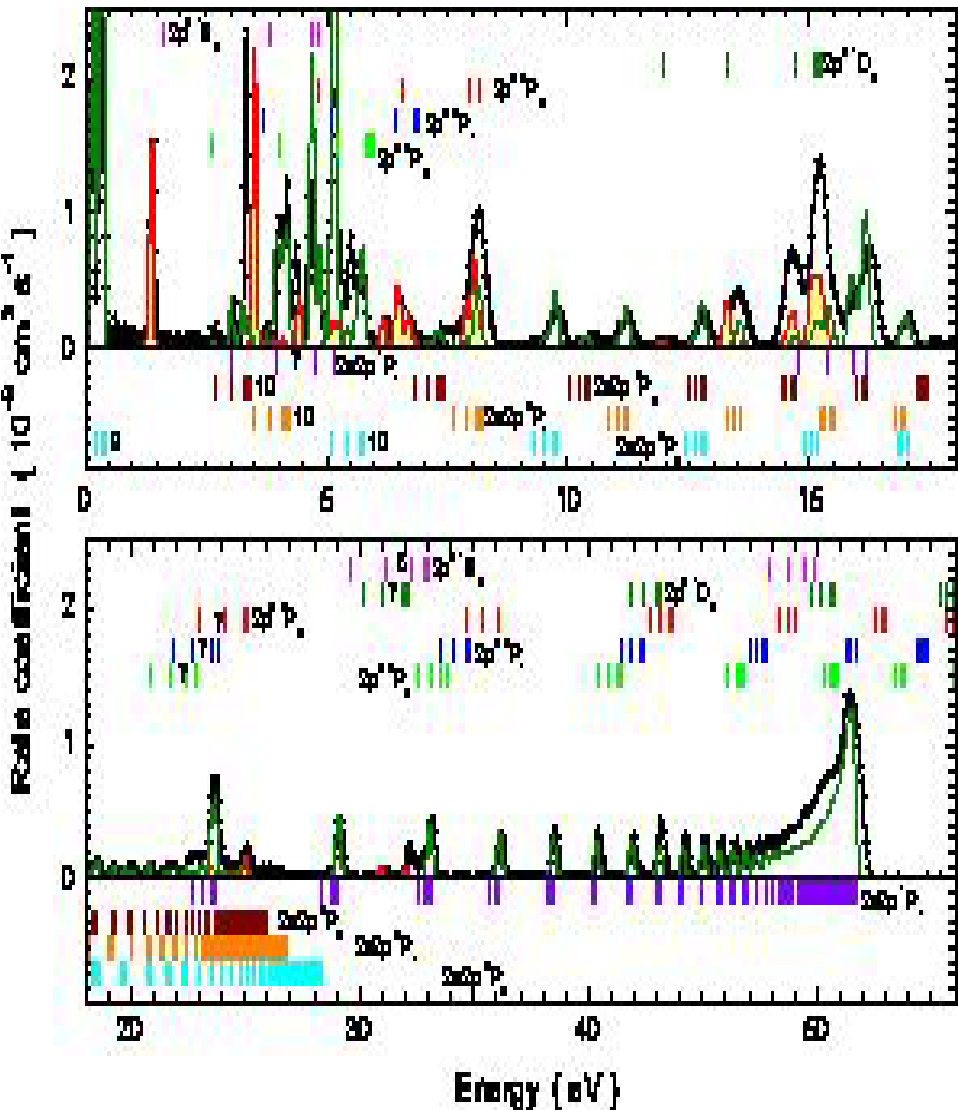
•  $\text{Fe}^{25+}$ : Ballance et al, J. Phys. B **35**, 1095 (2002).

# ***Dielectronic Recombination***

- Project well underway to calculate dielectronic recombination rate coefficients for complete isoelectronic sequences
- Currently completed H-like through Ne-like sequence;
- Moving onto second row of periodic table
- Theoretical approach is more accurate at high temperatures where recombination is more focused towards fusion plasma applications

- *Theory & Methodology: N. R. Badnell et al, A&A, accepted (2003).*
- *H-like: N. R. Badnell et al, A&A, submitted, (2003).*
- *He-like: N. R. Badnell et al, A&A submitted, (2003).*
- *Li-like: J. Colgan, M. S. Pindzola, and N. R. Badnell, A&A, submitted (2003).*
- *Be-like: J. Colgan, M. S. Pindzola, A. D. Whiteford, and N. R. Badnell, A&A, submitted (2003).*

# Dielectronic Recombination of $Cl^{13+}$

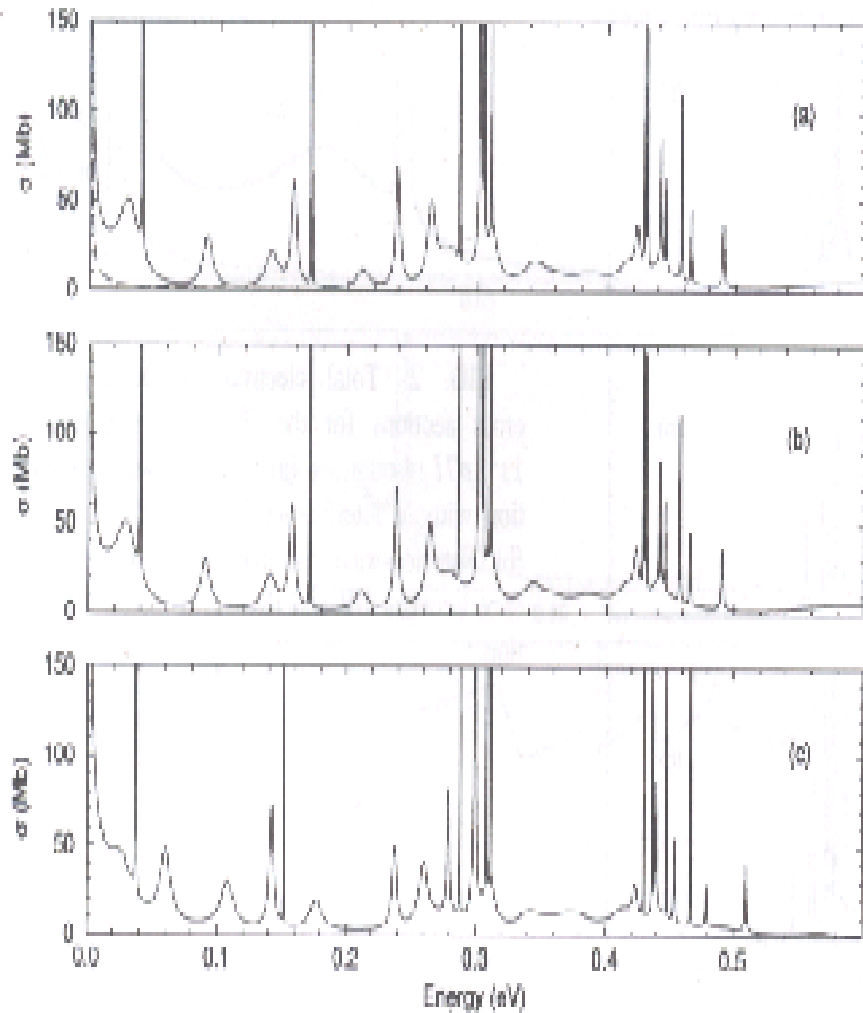


- Dielectronic recombination rate coefficients for Be-like  $Cl^{13+}$
- Comparisons of our theoretical calculations with experiment of Heidelberg group at the TSR
- Experimental confirmation of DR project results
- Trielectronic recombination (core excitation of two electrons) observed for the first time
- *M. Schnell et al. Phys. Rev. Lett. 88, XXXXXX (2003).*

# Problems for low temperature

## Dielectronic Recombination

$F^{6+}$



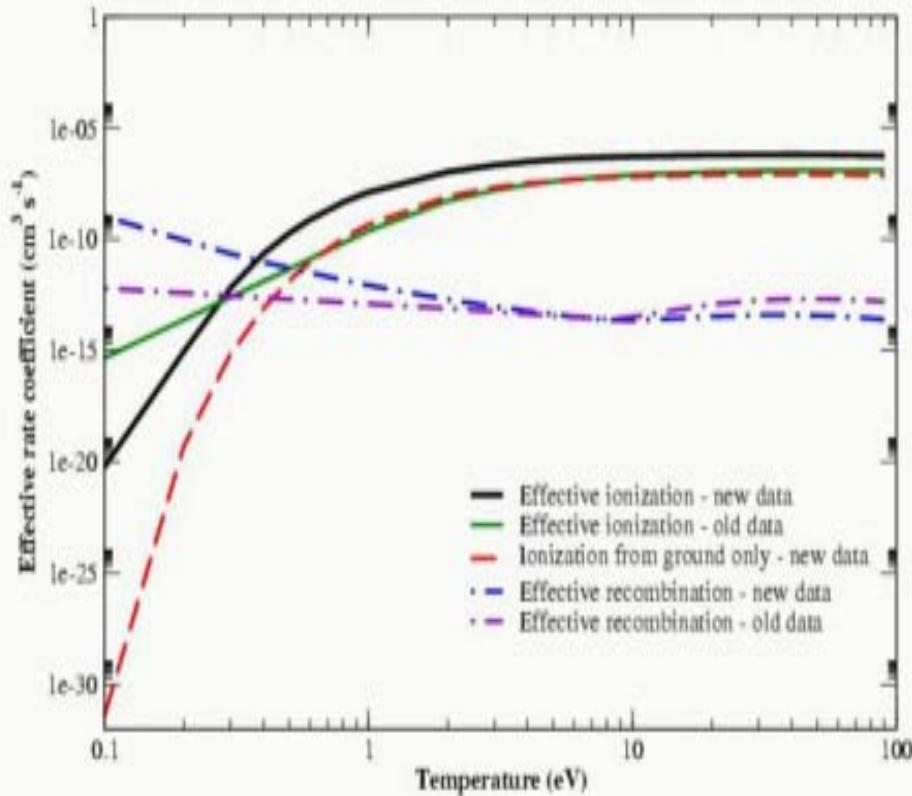
- a) DW calculations with no interference
- b) DW calculations with interference
- c) *R*-matrix calculation

- Very large structure calculations necessary to get position of resonances correctly
- Also need to include effects of interference between radiative and dielectronic recombination
- Also the effect of interacting resonances on recombination
- *Expt/theory for  $C^{3+}$ : S. Mannervik et al, Phys. Rev. Lett. 81, 313 (1998).*
- *Theory for  $F^{6+}$ : D. M. Mitnik, M. S. Pindzola, and N. R. Badnell, Phys. Rev. A 59, 3592 (1999).*
- *Expt/theory for  $F^{6+}$ : M. Tokman et al, Phys. Rev. A 66, 012703 (2002).*

# *Problems for high temperature Dielectronic Recombination*

- Near Rydberg limits weakly bound states can be strongly affected by electric and/or magnetic fields present. The plasma microfield and the plasma densities near threshold can also have a strong influence on the rates in these regions
- Calculations which take into account interacting resonances and DR in crossed electric and magnetic fields have been made:
  - *model calculation: F. Robicheaux and M. S. Pindzola, Phys. Rev. Lett. 79, 2237 (1997).*
  - *$C^{3+}$ ,  $Si^{11+}$ : F. Robicheaux, M. S. Pindzola, and D. C. Griffin, Phys. Rev. Lett. 80, 1402 (1998).*
  - *$C^{6+}$ ,  $F^{6+}$  expt: G. Gwinner et al, Phys. Rev. Lett. 84, 4822 (2000).*
  - *$O^{5+}$  expt/theory: S. Bohm et al, Phys. Rev. A 65, 052728 (2002).*
- Currently no generalized codes exist which take into account crossed field effects within a detailed structure calculation
- These huge calculations will require parallelization of existing recombination computer codes

# *Effective ionization rates for Li*



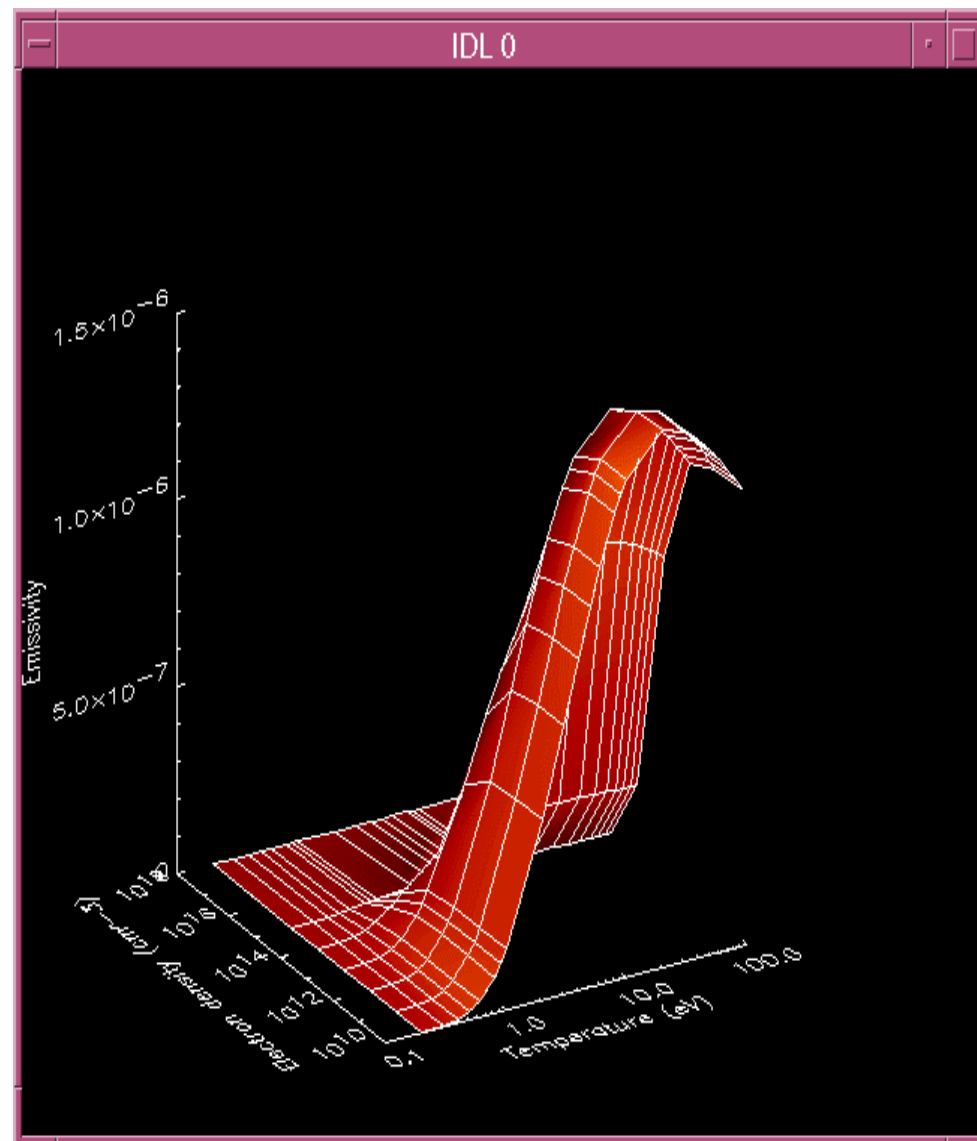
- New data for excitation, ionization, and recombination has been incorporated into an ADAS plasma modeling calculation
- Also available at CFADC database
- Very large differences compared to model using older atomic data
- Significant increase in the accuracy of the plasma modeling for lithium

*Effective ionization rate for Li at an electron density of  $10^{14} \text{ cm}^{-3}$  as a function of electron temperature*

- *S. D. Loch et al, Physica Scripta, submitted (2003).*

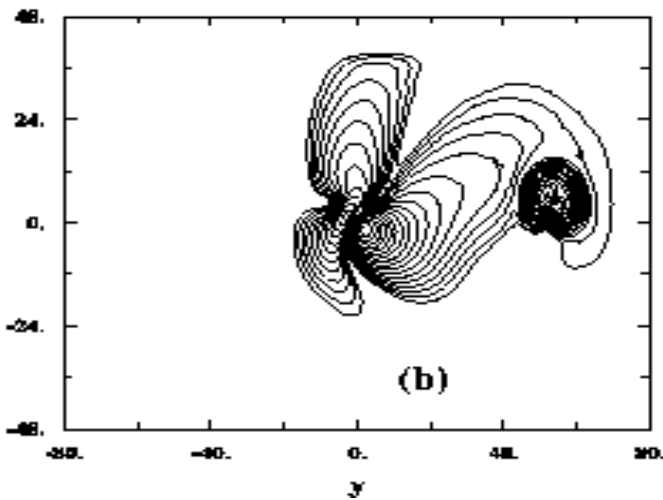
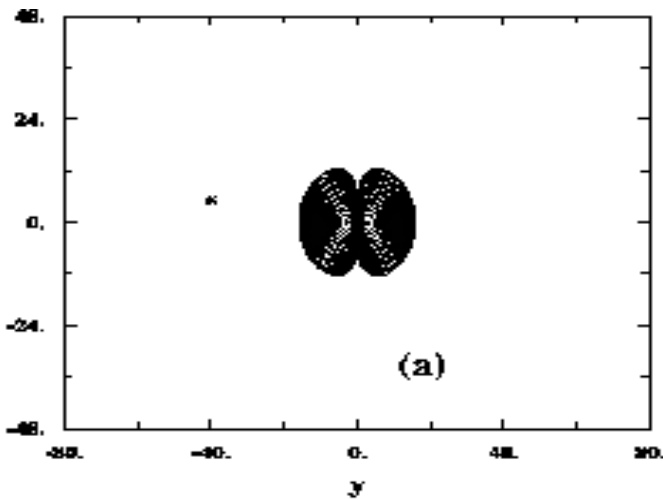


# *Li emissivity coefficients*



- Emissivity coefficients as a function of electron temperature and density
- Obtained from ADAS modeling calculation
- Data such as this is used in a wide range of fusion plasma edge diagnostics
- Also comparing these and similar plasma characteristics with modeling calculations made using Los Alamos set of computer codes
- *S. D. Loch, C. Fontes et al, Phys. Rev. E, submitted, (2003).*

# Ion-Atom Collisions

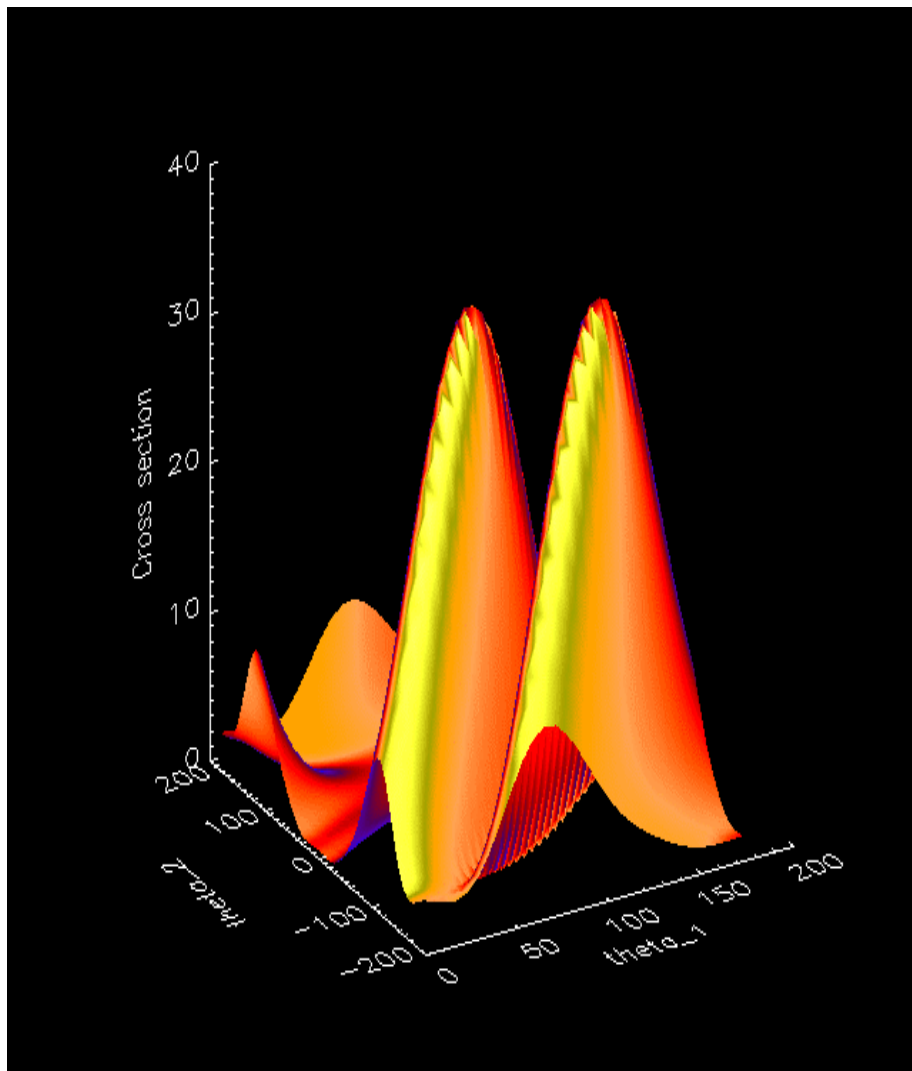


- $p + H$ : D. R. Schultz, et al, Phys. Rev. Lett. **76**, 2882 (1996).
- $p + H$ : J. C. Wells, et al, Phys. Rev. A **54**, 593 (1996).
- $p + H$ : A. Kolakowska, et al, Phys. Rev. A **58**, 2872 (1998).
- $p + H$ : D. R. Schultz, M. R. Strayer, and J. C. Wells, Phys. Rev. Lett. **82**, 3976 (1999).
- $p + Li$ : M. S. Pindzola, Phys. Rev. A **60**, 3764 (1999).
- $p + Li^*$ : M. S. Pindzola, Phys. Rev. A **66**, 032716 (2002).
- $\alpha + H$ : M. S. Pindzola, T. Minami, and D. R. Schultz, Phys. Rev. A **68**, XXXXXX (2003).
- Currently working on:

*Time-dependent evolution of  
electronic wave function during collision*

- $Be^{4+} + H$ :
- $p + H^*$ : in ExB fields

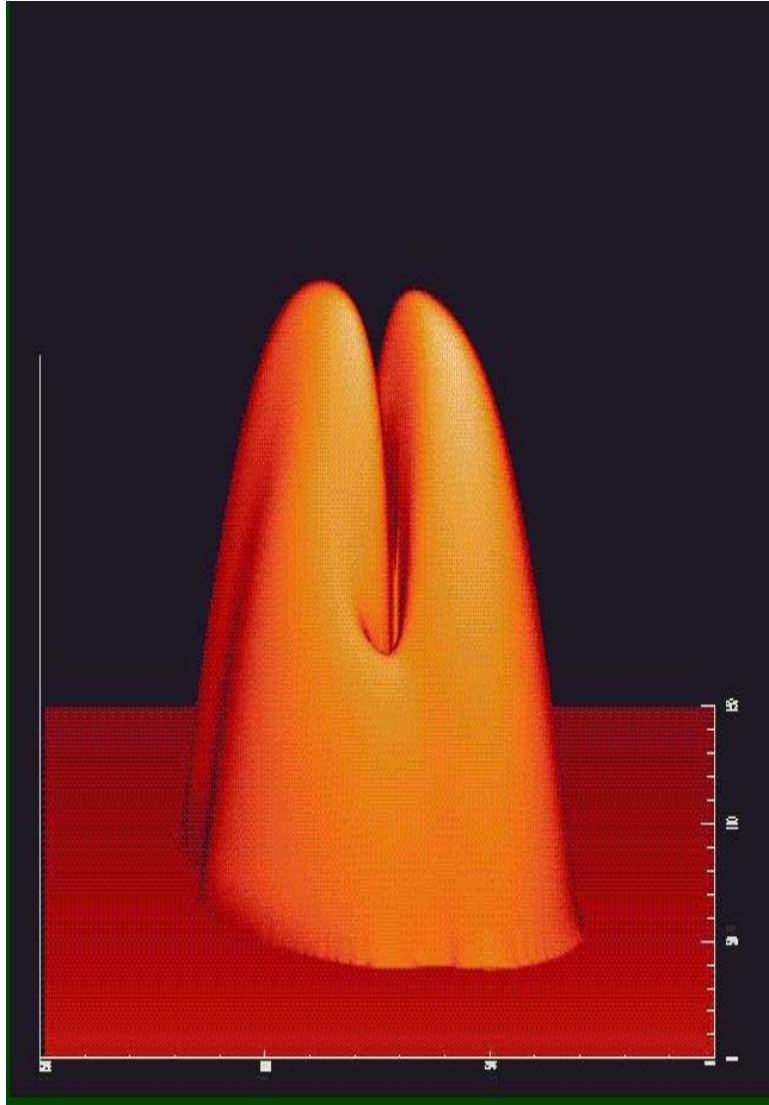
# *Double photoionization of helium*



- **$(\gamma, 2e)$  He:** J. Colgan, M. S. Pindzola, and F. Robicheaux, J. Phys. B **34**, L457 (2001).
- **$(\gamma, 2e)$  He:** J. Colgan and M. S. Pindzola, Phys. Rev. A **65**, 032729 (2002).
- **$(\gamma, 2e)$  Be:** J. Colgan and M. S. Pindzola, Phys. Rev. A **65**, 022709 (2002).
- **$(2\gamma, 2e)$  He:** J. Colgan and M. S. Pindzola, Phys. Rev. Lett. **88**, 173002 (2002)
- **$(\gamma, 2e)$  He\*:** J. Colgan and M. S. Pindzola, Phys. Rev. A **67**, 012711 (2003).

*Triple differential cross section at equal energy sharing between the ejected electrons*

# *Mode excitation of a BEC soliton state*



- D. L. Feder, M. S. Pindzola, L. A. Collins, B. I. Schneider, C. W. Clark, Phys. Rev. A **62**, 053606 (2000).
- J. Denschlag et al, Science **287**, 97 (2000).
- M. S. Pindzola and F. Texier, J. Phys. B **36**, XXX (2003).
- Currently working on BEC's in waveguides in collaboration with the MIT group of Pritchard and Ketterle.

*Density plot of a sodium condensate in a soliton state undergoing mode excitation*

# *Simulations of ultra-cold plasmas*

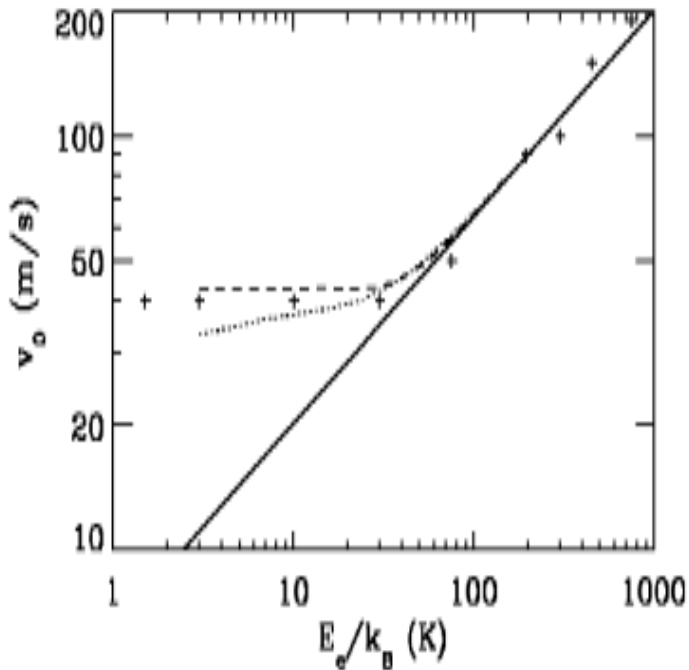


FIG. 2. Asymptotic expansion velocity,  $v_0$ , as defined in Ref. [2] as a function of the initial energy per ion. The crosses indicate representative experimental results [2]. The solid line is with no three-body recombination. The dotted line is with three-body recombination but no electron-Rydberg atom scattering. The dashed line is with three-body recombination and electron-Rydberg scattering. The calculations are for the initial average density of  $n_e = 10^{15} \text{ m}^{-3}$ .

- F. Robicheaux and J. D. Hanson, Phys. Rev. Lett. **88**, 055002 (2002).

- F. Robicheaux and J. D. Hanson, Phys. Plasmas **10**, June (2003).

*Inclusion of electron-Rydberg scattering and three-body recombination is necessary to properly simulate the expansion of an ultra-cold plasma*



## *E.g.: Supernova remnants (SNR's)*

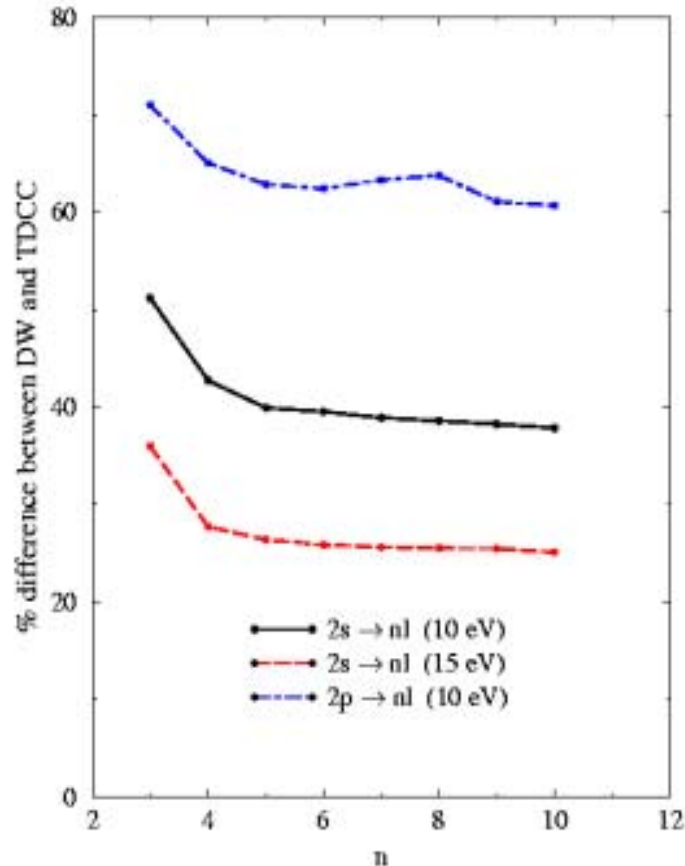
- Recent successful modeling of nebular spectra in Type Ia supernovae by including forbidden lines of cobalt ions
- Electron-impact excitation cross sections calculated using distorted-wave and  $R$ -matrix calculations for transitions in  $\text{Co}^{2+}$
- *W. Liu, D. J. Jeffery, D. R. Schultz, P. Quinet, J. Shaw, and M. S. Pindzola, Ap. J 489, L141 (1997).*
- *J. Shaw, M. S. Pindzola, N. R. Badnell, and D. C. Griffin, Phys. Rev. A 58, 2920 (1998).*



*The Crab Nebula – one of the earliest recorded Supernova explosions*

# *E.g.: Early Universe abundances of $^4\text{He}$ and $^7\text{Li}$*

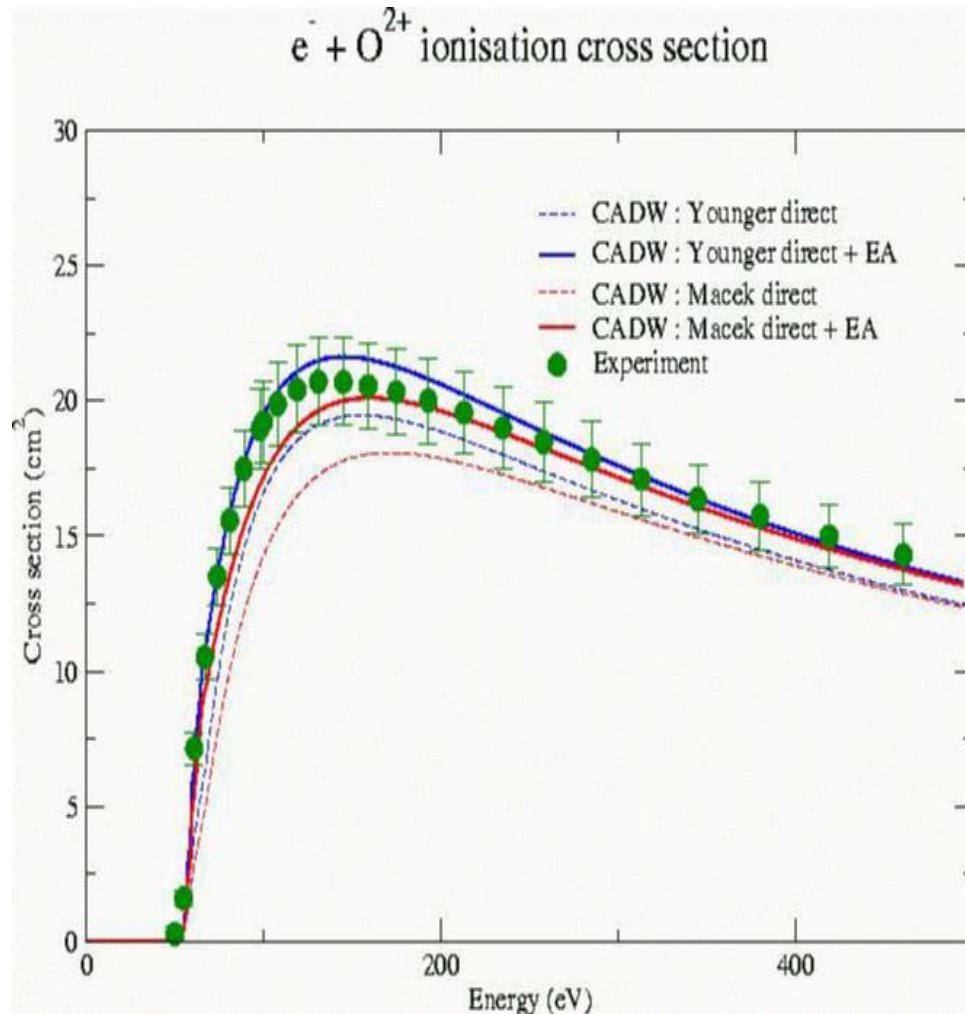
*Electron-impact excitation of neutral Li from 2s and 2p to high quantum number n*



- At low energies, TDCC method is around 50% different from DW calculations for excitation to all n-manifolds for Li and  $\text{He}^+$  systems
- Current databases use (at best) perturbative (DW) methods for these transitions – these may be in error by almost a factor of 2
- Working with G. Ferland to improve quality of He and Li atomic data used in astrophysical models
- *$\text{He}^+$* : M. C. Witthoeft, M. S. Pindzola, and J. Colgan, *Phys. Rev. A* 67, 032713 (2003).
- *Li*: M. C. Witthoeft, J. Colgan, and M. S.

## *E.g.: Cometary X-ray emissions*

- X-rays have been observed from many comets and their origin is still unclear
- Some unified plasma models explain this by modified two stream instabilities – Kellet et al
- However, they may also arise from radiative decays from solar wind ions following charge exchange with neutral species in the cometary atmosphere – Stancil et al
- Recent comparisons of ionization cross sections between experiment and theory for O ions can be used by modelers to increase accuracy of model



***S. D. Loch et al, Phys. Rev. A  
67, 042714 (2003).***

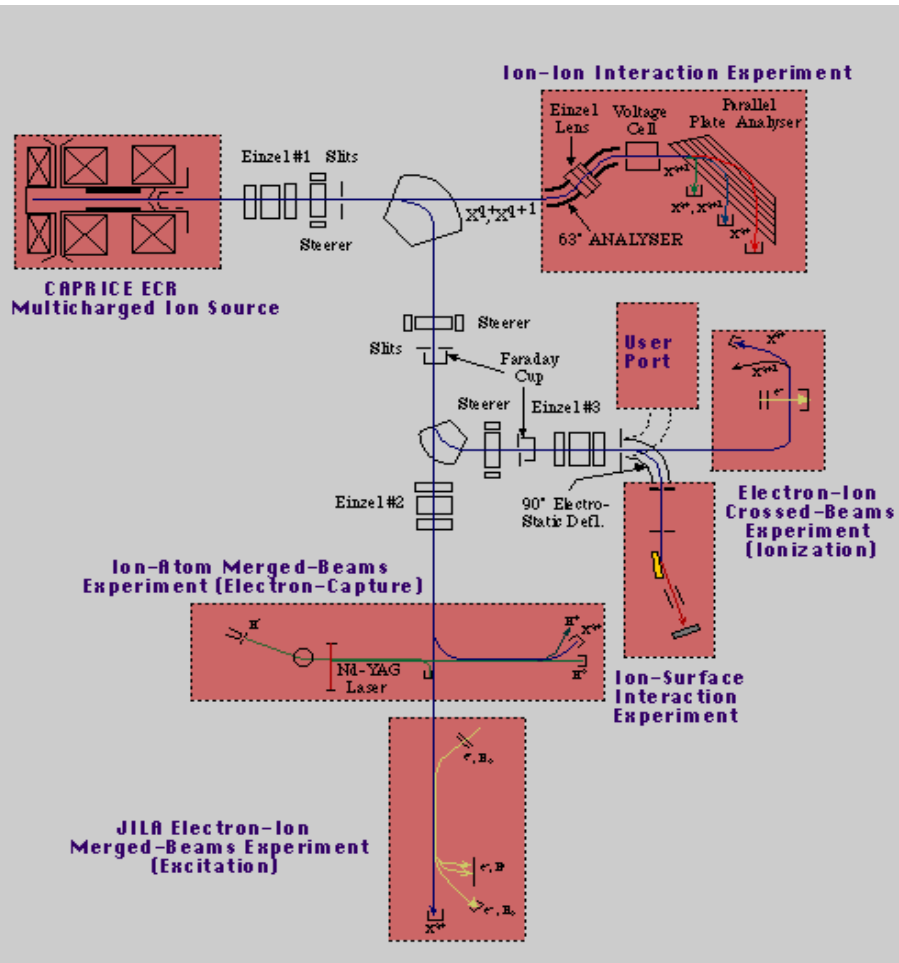


# *Planned calculations*

- Calculations of electron-impact ionization of ions in the Be-like, He-like and O-like isoelectronic sequences
- Many of these ions are directly observed by current X-ray observatories
- Ions in these sequences currently in astrophysical databases have significant uncertainties
- Theoretical calculations will support experiment; once theory and experiment agree, data will be used to produce more accurate ionization rate coefficients to be used in ionization balance calculations by astrophysical modelers
- Working closely with astrophysical modelers allows a guide to the systems which are important and selected for study

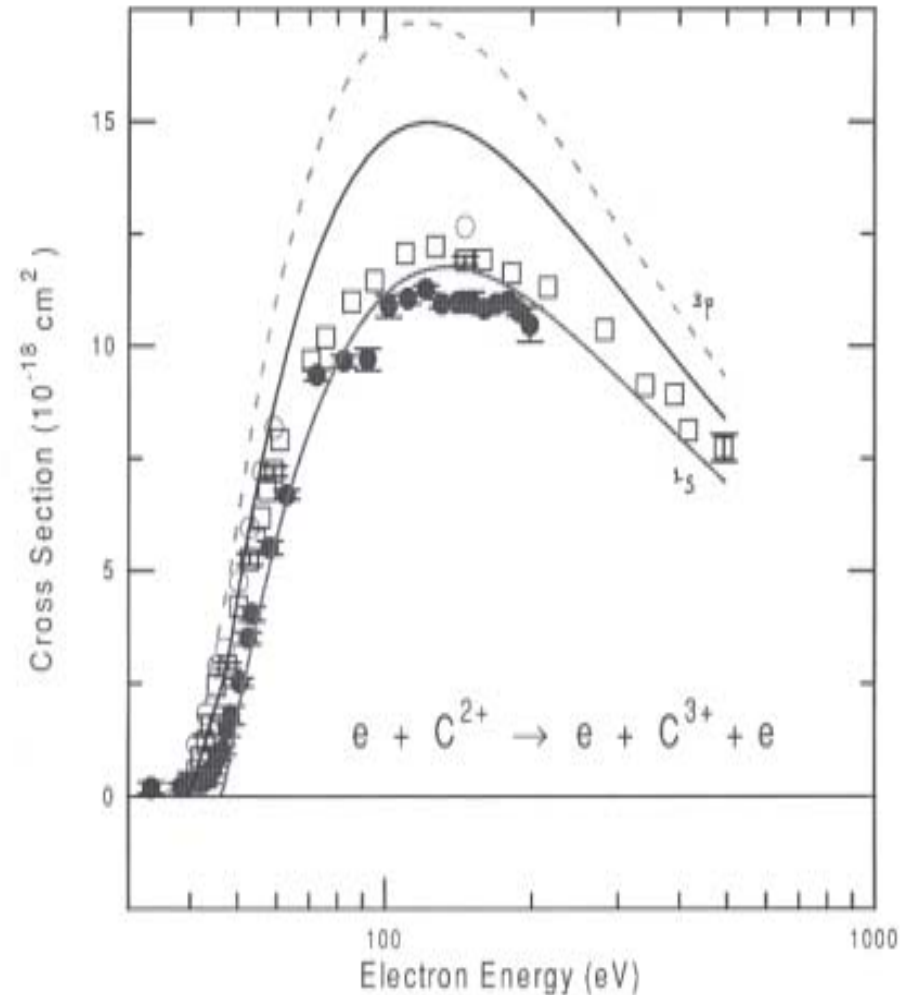
# *Electron-Impact Ionization Data for modeling cosmic plasmas*

- Detailed measurements are being planned at ORNL
- Unique facility in the U.S. as it measures both electron-impact excitation and ionization cross sections
- Allows measurements from metastable components of ion beams
- Major advance on previous measurements which could not give metastable component of ion beams - resulting in considerable uncertainties in the measurement



*Outline of the MIRROR (Multicharged  
Ion Research Facility) at ORNL, TN*

# *Future projects*



- Recent experiment of this process thought to be 40% ground ( $1S$ ) state and 60% metastable ( $3P$ )
- Configuration-average distorted-wave calculations up to 30% higher than experiment
- Unclear if due to uncertainties in experimental determination of metastable fraction; new experimental measurements planned
- New calculations planned using TDCC method – though significant theory development required

# *Future directions – electron collisions with molecular ions*

- Work in progress to calculate electron-impact excitation and ionization of light diatomic molecules
- Use time-dependent techniques; compare to  $R$ -matrix calculations
- Important in interactions of cosmic rays with atmospheric molecules
- Also diatomic molecules and molecular ions detected in interstellar clouds as well as in comets
- For many molecular species of astrophysical interest virtually no collisional data exists currently

# *Conclusions*

- Discussed range of atomic & molecular processes which are important in accurately modeling fusion and astrophysical plasmas
- A series of theoretical techniques have been developed and implemented by our group to try to address some of these atomic physics issues
- Our aim is to work closely with astrophysical plasma modelers in order to provide quality atomic data for systems of interest
- Close collaboration and good communication between groups will enhance progress that will be made!